#### **ORIGINAL ARTICLE**



### An updated review about carbon dioxide and climate change

Rex J. Fleming<sup>1</sup>

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#### Abstract

This manuscript will review the essence of the role of  $CO_2$  in the Earth's atmosphere. The logic of  $CO_2$  involvement in changing the climate will be investigated from every perspective: reviewing the historical data record, examining in further detail the twentieth-century data record, and evaluating the radiation role of  $CO_2$  in the atmosphere—calculating and integrating the Schwarzschild radiation equation with a full complement of  $CO_2$  absorption coefficients. A review of the new theory of climate change—due to the Sun's magnetic field interacting with cosmic rays, is provided. The application of this new theory is applied to climate-change events within the latter part of the Earth's interglacial period. The application to the Earth's Ice Ages is not detailed here due to manuscript size constraints, but is referenced for the reader. The results of this review point to the extreme value of  $CO_2$  to all life forms, but no role of  $CO_2$  in any significant change of the Earth's climate.

Keywords Carbon dioxide · Climate change · Solar magnetic field · Cosmic rays · Chaos

#### Introduction

The climate of planet Earth involves many nonlinear processes of the atmosphere, ocean, and other Earth sciences. This document will review what is known about these processes and explain an unfortunate misconception about *climate change*.

A specific definition of *climate change* is used to separate the term from the tremendous diversity of weather conditions that prevail on Earth. *Climate change* here implies the average surface temperature of the Earth adjusting upward or downward over a multi-year period (> 10-year). A significant external force is required to make such a change, e.g., a significant change in a solar property and/or a change in the Earth's albedo (the measure of how much the Sun's energy is reflected back into space).

Many believe and/or support the notion that the Earth's atmosphere is a "greenhouse" with  $CO_2$  as the primary "greenhouse" gas warming Earth. That this concept seems acceptable is understandable—the modern heating of the Earth's atmosphere began at the end of the Little Ice Age in 1850. The industrial revolution took hold about the same

Rex J. Fleming rex@rexfleming.com time. It would be natural to believe that these two events could be the reason for the rise in temperature. There is now a much clearer picture of an alternative reason for why the Earth's surface temperature has risen since 1850.

There is a thermal blanket or buffer for atmospheric surface conditions that have been in existence for the past billion years—existing in all "*climate-change regimes*" (warm or cold). Its exact form depends on the effective solar energy reaching the Earth's surface. The sources of the thermal blanket and the subsequent transfer of heat upward are from three forces: the Earth's gravitational field with its impact on convection, the condensation of water vapor (H<sub>2</sub>O), and the radiation effects of the two primary atmospheric trace gases of H<sub>2</sub>O and CO<sub>2</sub>. It will be demonstrated that the radiative roles of H<sub>2</sub>O and CO<sub>2</sub> are relatively minor, compared to their far more important role in maintaining sustained life on our planet.

The purpose here is to objectively examine every facet of the  $CO_2$ /climate-change issue. This will entail a review of the historical record of *climate change* and that of the twentieth-century warming relative to  $CO_2$  concentrations. The appropriate radiation code for the Earth's atmosphere will be examined in far more detail than is typically exposed in the scientific literature.

"The atmosphere and its extremely diverse weather" section describes the Earth's chaotic atmosphere and the diversity of the weather produced. This chaotic weather

<sup>&</sup>lt;sup>1</sup> Global Aerospace, LLC, 7225 Spring Drive, Boulder, CO 80303, USA

variability will continue to exist within each *climate-change* regime. "Does  $CO_2$  have a role in causing climate change? Observations examined" section reviews the early evolution of the atmosphere, and the ice ages of the past through the current modest ice age. Within ice ages, there are glacial periods with ice sheets and interglacial periods with little or no ice. Two different proposed reasons for the twentieth-century warming are examined.

Fourth section explains "The source of the Earth's thermal blanket". This daily recurring thermal protective layer, and the subsequent immediate heat transfer upward occur due to the three physical causes. These forces act to create a balance between the incoming solar shortwave and outgoing longwave energy.

"H<sub>2</sub>O and CO<sub>2</sub> in the radiation package" section quantifies the radiation roles of H<sub>2</sub>O and CO<sub>2</sub>. The standard approach to the radiation transfer is used—the integration of the Schwarzschild equation. "Solar magnetic field/cosmic ray factors affecting climate change" section summarizes the factors that cause *climate change* in the new theory of the interaction of the Sun's magnetic field with cosmic rays. "Summary" section provides a very brief review of the benefits of this important molecule, and a summary.

## The atmosphere and its extremely diverse weather

The atmosphere has a temperature profile which decreases with height through the troposphere which contains 80% of its mass. At the minimum temperature level of the troposphere is the tropopause—this varies with latitude and season. The tropopause height is generally 9 km over the poles and 16–17 km at the equator.

Above the tropopause is the stratosphere where temperature increases with height to about 50 km. The stratosphere obtains its heat by the direct absorption of the Sun's energy by ozone  $(O_3)$ .

The primary driver of Earth's weather systems is the very large-scale process of baroclinic instability that occurs independently in both hemispheres. Differential heating between the incoming solar radiation and the outgoing infrared radiation creates a pole-to equator temperature gradient and produces a growing supply of available potential energy. Eventually, the zonal thermal wind, developing to balance that temperature gradient, becomes baroclinically unstable. The resulting large-scale baroclinic waves transfer warm air poleward and cold air equatorward.

At the same time, the *eddy* (wave) available potential energy is converted into eddy kinetic energy by the vertical motion within the waves—maintaining the kinetic energy of the atmosphere against frictional dissipation. The waves intensify until the heat transferred poleward balances the radiation deficit. Various processes within the atmosphere (friction, radiation to space, etc.) damp the unstable waves and the baroclinic cycle is repeated.

Thompson (1987) developed a low-order general circulation model consisting of a single finite amplitude baroclinic wave interacting with the zonal mean shear flow, maintained against a friction parameter (**D**) and driven by a differential heating term (**H**)—thus containing all the requirements for baroclinic instability. The model produced accurate values for certain features including the *vacillation period* [the time from a zonal flow (non-wave) configuration to the maximum wave development, and back again to zonal flow] of approximately 23 days—close to that seen in the Southern Hemisphere (Webster and Keller 1975).

Vacillation, with fixed point attractors, were the only dynamic entities included and described in Thompson's original presentation. Later this model was found to produce two other attractor types: limit cycles and chaos (Fleming 2014). The chaos was produced over various values of the parameters **D** and **H**, and also produced by the process of sensitivity to the initial conditions—whereby on a strange attractor two initially close trajectories on the attractor eventually diverge from one another—exponentially over time.

The variability of the simple model above will be demonstrated, but this diversity would have been more intense with the 4–6 waves included (typically seen in a hemisphere) allowing wave–wave nonlinear interaction.

Figure 1 (Fleming 2014) indicates two solutions in the above model, vacillation, and chaos, respectively, where X(1) represents the mean horizontal temperature gradient and X(2) represents the net poleward heat transport. The longest vacillation cycle was 25.2 days. The longest chaos cycle was 35.0 days. The initial conditions were the same for all the parameters and variables—except that the initial mean horizontal temperature gradient [X(1)] was 67% of its fixed point value for the vacillation solution, and 66% of that value for the chaos solution.

Notice the significant difference in the scales of the two solutions in Fig. 1. The important difference between vacillation and chaos is that the range of X(1) and X(2) is approximately three times greater for the chaos case—the complete trajectory of the vacillation solution nearly fits inside the opening in the chaos trajectory. When such a large solution difference occurs from such two closely spaced initial points, this is explosive baroclinic instability (EBI).

A Monte Carlo approach evaluated the power of the chaos within this nonlinear model of baroclinic instability. Using a known chaos initial state for X(1), 40,000 different initial states were selected from a random number generator for a normal distribution with a standard deviation of only 0.001. The only model value changed for the 40,000 different deterministic runs was X(1)=0.4 + the normal deviate.

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**Fig. 1** Explosive baroclinic instability—vacillation on the left, chaos on the right (Fleming 2014)



The results for the 40,000 deterministic chaotic solutions were all different. The spread of solutions, using X(3) as an example (the wave kinetic energy considered a proxy for storm intensity), provided a dynamic range. The maximum value of X(3) within a chaos run was considered as a measure of the strength of that run. The average maximum X(3) for all the runs was 18.95. The minimum and maximum of this X(3) measure were 7.42 and 27.68—nearly a factor of four difference in magnitude.

This weather diversity would expand with the seasons with different heating characteristics. This weather variability would occur within any *climate-change* regime—warm or cold. However, there will never be runaway chaos as EBI is limited by the dynamics of the system (Fleming 2014).

Further changes in **H** and **D** were evaluated over a wide range. Large **H** and small **D** provide a fast system and small **H** and large **D** a sluggish system as anticipated. However, smaller values of **H** and **D**, as might be expected within an ice age, lean toward more chaos (Fleming 2015).

## Does CO<sub>2</sub> have a role in causing *climate change*? Observations examined

The Earth was formed 4.6 billion years ago, and the atmosphere 600 million years later. The peak of the CO<sub>2</sub> degassing rate coincided with the maximum tectonic activity about 2.7 billion years ago—reaching a value of approximately 10,000 times the current atmospheric value of 400 ppmv (parts per million by volume). Water degassing occurred much earlier, but reached its peak about 2.5 billion years ago. There then began a systematic (but oscillating) decrease in atmospheric CO<sub>2</sub> as it combined with water to provide the hydration of rock forming minerals in the oceanic and continental Earth crust (Sorokhtin et al. 2007).

A comparison of  $CO_2$  concentration in the atmosphere with some of the ice ages over time is presented from data

shown in Fig. 2 (Plimer 2009). Atmospheric  $CO_2$  continued to decrease (with oscillations up and down) until it had the value between 10 and 200 times today's concentration by 1.8 billion years ago (Kaufman and Xiao 2003).

Various Ice Ages are shown in Table 1 where the values of  $CO_2$  concentration lie between the two extremes. Clearly,  $CO_2$  values have no correlation with the ice ages.

Table 1 indicates  $CO_2$  concentrations for the ice ages and the warm periods between them—where the surface temperature was estimated to be ~8 °C higher than today's average value. The Earth's albedo apparently decreased during the warm periods—the Sun's insolation has not changed over the past billion years (Sorokhtin et al. 2007).

The table illustrates no correlation of  $CO_2$  and temperature. Why this is true must be looked at in detail and will be in "H<sub>2</sub>O and CO<sub>2</sub> in the radiation package" section.

Ice cores with sufficient vertical resolution (time resolution) have provided 420,000 years of data from Antarctica indicating that the temperature changes preceded the corresponding  $CO_2$  changes. An American team found the time lag (due to ocean mixing) of  $CO_2$  behind temperature of several hundred years. The oceanic reservoir of  $CO_2$  is far greater than that of the atmosphere. When the oceans are warm, they outgas  $CO_2$ , and when the oceans are cold atmospheric  $CO_2$  dissolves into the oceans (Fisher et al. 1999).

A subsequent study in 2003 by a French team indicating that deglaciation was not caused by  $CO_2$  which lagged the temperature by 200–800 years (Caillon et al. 2003). A third effort by Russian scientists arrived at the same conclusion, where the estimated delay was 500–600 years (Monin and Sonechkin 2005). This was claimed to be 420,000 years of data with undisputable evidence that  $CO_2$  concentrations of the atmosphere are the effect of global temperature changes and not their cause (Chilingar et al. 2008).

Now consider the *climate change* over the past 11,500 years of the Earth's interglacial period. There exists





| Table 1Ice ages andintermediate warm periods from65 to 850 million years ago | Various Ice Ages time (MYA)/name | CO <sub>2</sub> (ppmv) | Intermediate warm periods (MYA) $\Delta T = \sim 8 \ ^{\circ}\text{C}$ | CO <sub>2</sub> (ppmv) |  |
|--|----------------------------------|------------------------|--|------------------------|--|
|  | 850-630/Cryogenic                | ~40,000                | ~480–600   | 4200                   |  |
|  | 460–430/Andrean-Saharan          | 4000-4400              | ~360-420   | 3000                   |  |
|  | 350–260/Karoo                    | 370-400                | ~170–240   | 1200                   |  |
|  | 160–120/Scutum-Crux              | 2000–2400              | ~50–100  | 1000                   |  |

an excellent summary (Plimer 2009) of these alternating warm and cold periods. Only two of these are listed below.

Medieval Warming (900–1280) was a warm period where society thrived! The summers were long, the crops were plentiful. The population increased, cities grew, universities were established and cathedrals were built (Gimpel 1961).

The Little Ice Age (1300–1850) occurred in two major phases, with famine in Europe killing millions between 1690 and 1700 (Grove 1988). The Little Ice Age was initiated with what is called a "quiet Sun" a period of very few sunspots referred to as a *Solar Minimum*. The Maunder Minimum was the most important (1645–1715).

Cosmic rays are star dust—mostly hydrogen protons from exploding stars. These enter the Earth's atmosphere when the Sun is "quiet"—the solar wind and its magnetic field are weak.

An isotope is produced from the normal light element of beryllium (normally with 4 protons and 5 neutrons) into beryllium-10—it is produced by cosmic rays as follows. A cosmic ray entering the atmosphere creates a shower of secondary cosmic rays, e.g., an energetic neutron. This collides with an oxygen atom, removing a neutron for beryllium to make beryllium-10 (4 protons and 6 neutrons). Be-10 has a half-life of 1.4 million years.



Fig. 3 Be-10 values with solar activity—sunspot numbers from Hoyt and Schatten (1998), Be-10 from Beer et al. (1994)

Figure 3 indicates Be-10 values are strong when cosmic rays are strong (when the Sun is weak)—note the strong values during the important Maunder Minimum of the Little Ice Age. Be-10 values have weakened considerably with the strong Sun during the *Modern Warming*.

Another proxy for solar activity involves cosmic ray interaction with nitrogen (N-14, normally 7 protons and 7 neutrons). In this case, the energetic neutron collides with nitrogen and the atom loses a proton and gains a neutron



to become carbon-14 (6 protons and 8 neutrons). Normally the carbon atom is carbon-12 (6 protons and 6 neutrons). Figure 4 adapted from (Plimer 2009) indicates C-14 data are extremely small during the warm Medieval Warming and also in the Modern Warming since 1850. The C-14 values also have largest magnitudes at the sunspot minima.

The  $CO_2$  concentrations have relatively little change over the interglacial period (see Fig. 2). A Greenland ice core provided the value 270 ppmv dated from 600 years ago (Neftel et al. 1983). The pre-industrial revolution estimate by several methods provides a narrow range of 270–290. Further detail in "Solar magnetic field/cosmic ray factors affecting climate change" section indicates no correlation of  $CO_2$  with any of the climate changes observed and well documented during the 11,500 interglacial period.

Since the Little Ice Age, a strong Sun is revealed by both Be-10 and C-14 decreases. The total magnetic flux leaving the Sun (dragged out by the solar wind) has risen by a factor of 2.3 since 1901 (Lockwood et al. 1999). The strong solar magnetic field has shielded the Earth from cosmic rays and is the cause of the Modern Warming that has occurred through to the current time. This competing climate theory is from Svensmark and Friis-Christensen (1997).

The theory involves the interaction of the solar magnetic field with cosmic rays. When the solar magnetic field is strong, it acts as a barrier to cosmic rays entering the Earth's atmosphere, clouds decrease and the Earth warms. Conversely when the solar magnetic field is weak, there is no barrier to cosmic rays—they greatly increase large areas of low-level clouds, increasing the Earth's albedo and the planet cools. The solar magnetic field is generated by the solar dynamo with the principal cause being the angular momentum of the Sun's differential rotation (Charbonneau 2014). The Sun's equatorial region rotates faster—24 days, compared to the polar regions which rotate once in ~ 30 days. This solar dynamo accounts for the variability of the sunspot amplitudes and frequency changes.

There is another factor affecting the solar dynamo that occurs on a longer time scale. This is the Sun's motion about the center of mass of the solar system—the solar system barycenter (SSB). The position of the SSB is constantly changing primarily as a function of the mass of the Sun and the

Atmospheric CO<sub>2</sub> at Mauna Loa Observatory



Fig. 5 CO<sub>2</sub> record from Mauna Loa from NOAA/CO<sub>2</sub> data/full record

four major planets (Sharpe 2008). The Sun's travel about the SSB adds its orbital angular momentum from that journey to its own rotational angular momentum so that both contribute to important changes in the Sun's magnetic field intensity. Details of all the factors that influence this solar magnetic field/cosmic ray interaction are described with references in "Solar magnetic field/cosmic ray factors affecting climate change" section.

This new climate-change theory competes with the  $CO_2$  warming associated with the timing of the industrial revolution. The combination of both of these two solar magnetic field influences appears to be the cause of a twentieth-century cooling within the Modern Warming. Thus, it is imperative to consider the twentieth-century temperature record since the industrial revolution.

The World War II and postwar period was a time of tremendous industrial growth from 1940 to 1975 (Plimer 2009). Figure 5 indicates the modern  $CO_2$  record from Mauna Loa which indicates the increased activity. However, there was a significant drop in temperature from coastal stations around the Arctic Ocean from 1940 to 1970 of 1.4 °C (Solomon



Fig. 6 Near-surface air temperature change in the twentieth century (see data sources in box within the figure)

2008). Also Fig. 6 shows a global cooling between 1940 and 1975 of surface temperatures over land (90 N–60 S) from three different records. The benchmark indicator for the  $CO_2$  warming theory, the NOAA Mauna Loa carbon dioxide record, fails to indicate this 35-year period of cooling. There are two solar-related events that support this cool climate–change event.

In contrast to the smoothed decrease in the Be-10 data in Fig. 3 from the post Little Ice Age—a closer look at the background shows sharp spikes of significant Be-10 increases from 1900 to 1970 indicating increased cosmic rays. A second solar record below provides further evidence of this cool period.

The sunspot cycle has an average period of 11.2 years, but the length varies from 8 to 14 years. The length of a sunspot cycle (LSC) is an indicator of the Sun's eruptional activity. The Gleissberg (1965) cycle resulted from his smoothing of the time series of the length of the sunspot cycles (LSC) and a secular cycle of 80–90 years emerged.

Figure 7 is from Landscheidt (2003) where Gleissberg's smoothed data were displayed. The heavy line is the smoothed LSC line and the light line is the land air temperature in the Northern Hemisphere. The heavy line agrees very well with the temperature and also with the temperature record of Fig. 6 with the cooling from 1940 to 1975. It appears that the atmospheric temperature is oblivious to  $CO_2$ concentration! Why? One must check the radiation calculations in "H<sub>2</sub>O and CO<sub>2</sub> in the radiation package" section to see whether there is something special about the role of CO<sub>2</sub>.



Fig. 7 Gleissberg record of LSC found in Landscheidt (2003)

#### The source of the Earth's thermal blanket

The Sun's output has been steady for the last one billion years. Three processes have maintained the energy balance over that time period—any systematic deviation in either direction over such a long period of time would have made life on Earth impossible—a planet too hot (a burnt cinder) or too cold (a ball of ice).

These three processes are: radiation, latent heat release from condensation, and atmospheric convection. These all work together to heat the lower atmosphere, transfer heat upward and cool the upper atmosphere, and achieve radiative balance.

The Earth receives incoming solar radiation and radiates terrestrial longwave radiation. This radiant energy travels in waves at the speed of light  $(3 \times 10^8 \text{ m/s in a vacuum})$ . Details of radiation are described in "H<sub>2</sub>O and CO<sub>2</sub> in the radiation package" section.

Water exists on Earth in three phases—solid, liquid and gas. The intermolecular forces in water molecules are decreased as energy (heat) is applied to the phases of water. The most important phase changes of water for the climate system are the Sun's energy (2500 J/g) evaporating water from the oceans, and then that energy being released as latent heat of condensation.

Convection is clearly the most important mechanism for transferring heat upward (Emanuel 1994). As hot air expands it becomes less dense and rises. Similarly, denser cooler air drops down and replaces the warmer air. This is a diabatic process as there is always some entrainment (conduction—molecular collisions exchanging heat—a minor process within a gas, but nevertheless present.)

Sorokhtin et al. (2007) have provided a way to quantify these processes in relative terms. Their procedure is clever and provides representative values for the three forces—at least in a space/time averaged sense. The values for the three processes are:

diabatic convection: 0.2394/0.3597 = 66.56%, diabatic condensation of water vapor = 0.0896/0.3597 = 24.90%, diabatic radiation (primarily H<sub>2</sub>O and CO<sub>2</sub>) = 0.0307/0.3597 = 8.54%.

These values are useful, but should not be taken too seriously—they are, at best, ensemble averages over space/time. Their sum may be close to 100%, but there can be significant variability among the processes. This diversity is assured by the chaotic nature of the atmosphere. The radiation component has the smallest percentage impact, but has an important day-to-day role in providing energy balance.

#### $H_2O$ and $CO_2$ in the radiation package

Figure 8 is a visual of the spectrum of longwave radiation adsorbed and emitted from the Earth and its atmosphere. The abscissa represents wavelengths of radiation in microns



Fig. 8 Absorption spectrum for H<sub>2</sub>O and CO<sub>2</sub> (Rothman et al. 2009)

 $(10^{-6} \text{ m or } 10^{-4} \text{ cm and "microns" are shortened to } \mu\text{m})$ . Incoming solar energy for climate is important in the short range of  $0.1-2.0 \mu\text{m}$ —and for longwave terrestrial radiation the range is from 4.0 to 40  $\mu\text{m}$  (Peixoto and Oort 1992).

Radiation interacts with matter on both the atomic and molecular level. Gases in atomic form adsorb and emit radiant energy in very narrow wavelengths that result from quantized changes in electronic states—called spectral absorption lines. Vibrational absorption occurs within a molecule due to the vibration of component atoms about their mean position within the molecule. Rotational absorption is due to the rotation of a molecule around its center of mass. The multiplicity of vibrational–rotational modes creates a complex irregular absorption spectrum with bands containing thousands of lines. The strongest of the H<sub>2</sub>O lines have not the strength of the strongest of the CO<sub>2</sub> coefficients.

One can make a broad brush comparison of the relative roles of  $H_2O$  and  $CO_2$  in the heating of the thermal blanket. The units on the coefficients in Table 2 are in m<sup>2</sup>/kg. The comparison is for the level of the thermal blanket considered to be one km thick within the planetary boundary layer.

The concentration of  $CO_2$  is considered to be uniform over the atmosphere at 400 ppmv. The concentration of water vapor varies from a maximum of 40,000 ppmv (Hong Kong) to the lowest measured value of 4 ppmv in the upper stratosphere. A value for water vapor at one km is estimated to be 11,000 ppmv, so the ratio of mass of H<sub>2</sub>O/CO<sub>2</sub> at one km is approximately 11,000/400=27.5. Comparison of the absorption coefficients over the full range of 1.5–18 µm gave the result:  $CO_2/H_2O = -5.5$ . Thus, water vapor dominates by the ratio of 27.5/5.5=5.

The CO<sub>2</sub> absorption coefficients obtained directly from the Pacific Northwest National Laboratory (Rothman et al. 2009) were 3,900,000 in number ranging from 1 to 40  $\mu$ m. Based on the author's limited personal computer/software, only every 10th line was used bring the total down to 390,000. Table 3 indicates the values of those coefficients according to magnitude. The reduction was implemented so that the true maximum coefficient was used.

These two data sets are statistically equivalent. Table 3 indicates that the number of coefficients considered transparent K < 1 is 98%—and the maximum is the same in both data sets. The average value of each is virtually the same. In all

| Table 2 Absorption coefficients                       |
|---|
| for H <sub>2</sub> O and CO <sub>2</sub> from Pacific |
| Northwest National Laboratory/                        |
| Rothman et al. (2009)                                 |

| Band name   | Range wavelength $(\mu m)$ | Max value (m <sup>2</sup> /kg) | Average value    |
|---|----------------------------|--------------------------------|------------------|
| H <sub>2</sub> O Band 1                           | 2.55–2.84                  | 78.02 at 2.6705 µm             | 2.10             |
| H <sub>2</sub> O Band 2                           | 5.00-7.10                  | 82.83 at 5.9351 µm             | 2.10             |
| H <sub>2</sub> O window<br>CO <sub>2</sub> window | 8.0–16.0<br>5.0–13.0       | 1.2<br>0.06                    | 0.0063<br>0.0005 |
| CO <sub>2</sub> Band 1                            | 4.20–4.50                  | 4596 at 4.2346 µm              | 68.4             |
| CO <sub>2</sub> Band 2                            | 13.61–16.00                | 596.1 at 14.98 µm              | 8.8              |

| Table 3 | $CO_2$ absor | ption coefficients | from PNNL |
|---------|--------------|--------------------|-----------|
|---------|--------------|--------------------|-----------|

| CO <sub>2</sub> absorption coefficients<br>(K) |         | 390,000 lines     | 3,900,000 lines   |  |
|--|---------|-------------------|-------------------|--|
| Units are m <sup>2</sup> /kg                   | 1–40 µm | K < 1 transparent | K < 1: transparen |  |
| %Transparent                                   |         | 98.00             | 98.00             |  |
| Maximum value                                  |         | 4596              | 4596              |  |
| Average value                                  |         | 1.2503            | 1.2482            |  |
| K<0.0001                                       |         | 228,424           | 2,284,198         |  |
| $0.0001 \ge K < 0.001$                         |         | 102,982           | 1,029,996         |  |
| $0.001 \ge K < 0.01$                           |         | 30,004            | 300,007           |  |
| $0.01 \ge K < 0.1$                             |         | 11,994            | 119,933           |  |
| $0.1 \ge K < 1.0$                              |         | 8811              | 88,201            |  |
| $1.0 \ge K < 10.0$                             |         | 4438              | 44,289            |  |
| $10.0 \ge K < 100.0$                           |         | 2695              | 26,853            |  |
| $100.0 \ge K < 1000$                           |         | 544               | 5453              |  |
| K>1000   |         | 108               | 1070              |  |
| Total lines                                    |         | 390,000           | 3,900,000         |  |

categories of the various magnitudes of the coefficients, the

number in the data set with 390,000 coefficients is approxi-

equation for the intensity of radiation, and the integration of

the Schwarzschild equation for net diffuse radiation. One

can see Houghton (1985) and Liou (2002) for details. The

The important equations for radiative transfer are Planck's

mately 10% of the number in the larger data set.

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Fig. 9 Radiative power intensity expressed by Planck's equation (this plot by Hashemi 2010)

 $dI_{\lambda}/k_{\lambda}\rho ds = -I_{\lambda} + B_{\lambda}(T)$ ; where the first term on the righthand side denotes the reduction in radiant intensity due to absorption, whereas the second term represents the increase in radiant intensity arising from the blackbody emission of the material.

The atmosphere is considered to be in thermodynamic equilibrium and is plane parallel. A differential optical depth can be defined as  $d\tau = -k_{\lambda}(z)\rho(z)dz$ . The coordinates are below from Liou.

| 0              | Z∞ | <br>T∞ | u1 | 0  |
|----------------|----|--------|----|----|
| τ              | z  | <br>ť  | ú  | P  |
| τ              | Z  | <br>т  | u  | Ρ  |
| τ              | z  | <br>т́ | ú  | P  |
| τ <sub>s</sub> | 0  | <br>Τs | 0  | Ps |

equation is found in radiation texts and the curves below computed for values of wavelength ( $\lambda$ ) and *T*.

The integration of the Schwarzschild equation is used in the computations below. The largest impact from level to level in this equation is the change in the intensity of the Planck equation (see Fig. 9) and note how it changes with temperature and wavelength). The formula is:

$$B(\lambda, T) = \left[2\Pi hc^2/\lambda^5\right] \left[\exp\left(ch/k\lambda T\right) - 1\right]^{-1}$$

Consider the transfer of thermal infrared radiation emitted from the Earth and the atmosphere where a beam of intensity will undergo the *absorption* and *emission* processes simultaneously. The Schwarzschild equation for this process is: The coordinate systems in  $\tau$ , Z, u, T and P for IR radiative transfer are shown above. The path length (u) is for absorbing and emitting gases (they absorb and warm and emit and cool) defined for the surface upward. The total path length is defined as  $u_1$ .  $T_{\infty}$  and  $Z_{\infty}$  are temperature and height at the top of the atmosphere. The surface temperature =  $T_S$ . The surface pressure is  $P_S$ . Z is a reference level.

The Schwarzschild equation solution using the Liou notation above and equations of Houghton is:  $F = -\int B(\lambda, T)(d\tau^*)/du)du + \int B(\lambda, T)(d\tau^*)/du)du$ ; the optical depth  $(d\tau^*)$  is due to the radiation being diffuse rather than a parallel beam.

In the first integral, the integration proceeds *downward* along the optical path from the reference level (Z) with

optical path *u* downward to the surface where the optical path = 0. In the second integral, the integration proceeds upward from the reference level (Z) with optical path u upward to the top of the atmosphere where the total optical path is  $u_1$ . Both integrals are positive, because of the convention that the path length is measured positive down and then positive up, respectively. The net flux at a level is the upward flux at the bottom of a layer minus the downward flux at the top of a layer.

Calculations begin with the CO<sub>2</sub> lines and coefficients in Band 1. The maximum coefficient in this band is quite large—the value of this coefficient  $K_{\lambda} = 4596 \text{ m}^2/\text{kg}$  and occurs at the wavelength of  $\lambda = 4.2346463 \mu\text{m}$ . All the 390,000 lines are organized by increasing wavelength, paired with their surface absorbance coefficient. One starts with a formula for the lines. The number of lines displayed in this presentation will vary from 70,000 to 300,000. Many more runs were performed over a wide range of lines—all of these runs provided similar numbers with the same conclusion. Band 1 run uses the formula below with J = 1-70,001.

 $\lambda = 1.0 + (J-1) * 0.00005 + 0.0346463$ . This provides lines from 1.0 to 4.50346463 µm. The value of J = 64,001provides a direct hit on the largest absorption coefficient at  $\lambda = 4.2346463$ . Every line in the formula is evaluated at every level ( $\Delta$ height = 1 km) with the following steps for each line and level. Information for each level is saved going upward, then downward.

- (1) The line is selected from the formula for J = 1-70,001. Note that the only line guaranteed to be exact is  $\lambda = 4.2346463$ . Other lines may also be exact, but in any case they are extremely close and linear interpolation provides the proper coefficient.
- (2) The standard temperature *T* at that height is selected for use in the Planck subroutine along with the wavelength

to arrive at the proper Plank radiation intensity for that  $B(\lambda, T)$ .

- (3) Rather than separate each reduction with height—i.e., the reduction in the Planck function B<sub>λ</sub>(λ, T), the reduction due to changes in the CO<sub>2</sub> density; and the reduction to the coefficient magnitude with temperature, one can obtain the same final answer by incorporating those changes into a single coefficient of reduced intensity (K<sub>E</sub>). The first of the three steps is the Planck change with height: K<sub>E</sub> = K<sub>λ</sub> × [B(λ, T)/B(λ, T<sub>Surface</sub>)] [note the original surface K<sub>λ</sub> is used for the first level (1 km)].
- (4) The new  $K_{\rm E}$  is further reduced by the density change in  ${\rm CO}_2$ .

 $K_{\rm E} = K_{\rm E} \times \rho(T, P) / \rho (T, P)_{\rm Surface}$  where  $\rho = P/1.889 T$ . [Where  $P = \rho R T$ , and R is for CO<sub>2</sub>.]

- (5) The new  $K_{\rm E}$  is further reduced by the decreasing temperature with height as  $K_{\lambda}$  increases in line strength with height.  $K_{\rm E} = K_{\rm E} \times (T/T_{\rm Surface})$ .
- (6) The final step is to correct for the path length since the radiation is diffuse. Several options, see Houghton (1985) and Liou (2002), all produce similar results. The options ranged from (0.602 reduction for strong coefficients and 0.5 for weak) to 0.6 for all coefficients. Both Bands 1 and 2 calculations used  $K_{\rm E} = 0.6 K_{\rm E}$  for all the absorption lines.

Having performed all the steps above, the data are saved, checked and statistics determined for each level: including the percent transparent determined for all  $K_{\rm E} < 1$ . Note that K in all the tables is  $K_{\lambda}$  at the surface and  $K_{\rm E}$  for all other layers above the surface.

These results for Band 1 (Table 4) reveal several important points. The first is that 93.68% of the surface coefficients  $K_{\lambda}$  are transparent with values < 1—quite a large number. This is important as all the CO<sub>2</sub> molecules are influenced by all the coefficients. The derived coefficients of reduced

| Lines Schwarzschild    | 70,001  | 70,001 | 70,001 | 70,001 | 70,001  | 70,001     |
|------------------------|---------|--------|--------|--------|---------|------------|
| K = coefficient        | Surface | 5 km   | 8 km   | 9 km   | 12 km   | 17 km      |
| %Transparent           | 93.68   | 99.37  | 99.95  | 100    | 100     | 100        |
| Max K                  | 4596    | 60.582 | 5.104  | 0.6902 | 0.0240  | 0.0148     |
| Average K              | 5.722   | 0.068  | 0.0124 | 0.0008 | 0.00003 | 0.00000005 |
| K<0.0001               | 26,410  | 58,947 | 64,542 | 65,619 | 68,328  | 70,001     |
| $0.0001 \ge K < 0.001$ | 22,089  | 4352   | 1927   | 1589   | 1358    | 0          |
| $0.001 \ge K < 0.01$   | 9236    | 2639   | 2093   | 2088   | 264     | 0          |
| $0.01 \ge K < 0.1$     | 4567    | 1832   | 1160   | 568    | 51      | 0          |
| $0.1 \ge K < 1$        | 3273    | 1787   | 243    | 137    | 0       | 0          |
| $1 \ge K < 10$         | 1915    | 309    | 36     | 0      | 0       | 0          |
| $10 \ge K < 100$       | 2002    | 135    | 0      | 0      | 0       | 0          |
| $100 \ge K < 1000$     | 407     | 0      | 0      | 0      | 0       | 0          |
| K>1000                 | 102     | 0      | 0      | 0      | 0       | 0          |
|                        |         |        |        |        |         |            |

intensity are 100% transparent from 9 to 17 km. Despite the very powerful absorption coefficients in Band 1 (there were 102 values greater than 1000), the influence of the Planck function is very strong.

The influence of line shape (see references) was included in the above calculation. These are important only for *strong lines* and they were included only when  $K_E$  was > 50. Removing the line shape changed the numbers slightly, but not the result. Different diffuse radiation options were used and provided the same primary result—each option provided the same levels having 100% transparent values.

The most important difference for Band 2 is the longer wavelengths where the Planck differential with height is considerably less—as indicated in Table 5. Many runs were made over the Band 2. All the runs with various numbers of lines and over various ranges of lines gave essentially the same bottom line results. Thus, the decision was made to minimize space and display runs of simultaneously calculating Bands 1 and 2 for all wavelengths between 1 and 30  $\mu$ m. Two runs were performed with 150,000 and 300,000 lines.

The temperature profile used for these calculations was that for the standard atmosphere, however, a very wide range of profiles scattered about that standard (some very stable, some very unstable) also produced the same upper-level transparency—all occurring first at 9 km.

These two Schwarzschild runs are shown together in Table 6—both runs give essentially the same results with 100% transparency at 9 km through 17 km. The maximum coefficient at 9 km is 0.69 in both cases. The CO<sub>2</sub> coefficients had > 95% of the surface coefficients already transparent. The maximum coefficient of reduced intensity  $K_{\rm E}$ =0.69 at 9 km implies only a slight residual of heat available at that altitude—though the coefficient is virtually transparent. All the remaining smaller coefficients are transparent—also allowing heat to space un-impeded.

The volume of  $H_2O$  at the one km level alone is capable of absorbing all the available solar heat at the surface, and does absorb five times that of  $CO_2$ . All the heat adsorbed at the surface was fully redistributed vertically by all the molecules with the help of all the coefficients.

| <b>Table 5</b> Planck intensity: $B(\lambda, T)$ for the strongest coefficients in Bands 1 and 2 | Height (km) | Temperature | $B(\lambda, T)/B(\lambda, T_0)$<br>$\lambda = 4.23466$ | $B(\lambda, T)/B(\lambda, T_0)$<br>$\lambda = 14.9815$ | $CO_2$ density<br>$\rho(T)/\rho(T_0)$ | $T/T_0$ |
|--|-------------|-------------|--|--|---------------------------------------|---------|
|  | 0           | 288.15      | 1.0000   | 1.0000   | 1.0000                                | 1.0000  |
|  | 1           | 281.65      | 0.7619   | 0.9235   | 0.9075                                | 0.9774  |
|  | 3           | 268.65      | 0.4252   | 0.7791   | 0.7422                                | 0.9323  |
|  | 5           | 255.76      | 0.2236   | 0.6466   | 0.6010                                | 0.8872  |
|  | 7           | 242.65      | 0.1098   | 0.5265   | 0.4813                                | 0.8421  |
|  | 9           | 229.65      | 0.0497   | 0.4192   | 0.3807                                | 0.7970  |
|  | 11          | 216.15      | 0.0205   | 0.3251   | 0.2971                                | 0.7519  |
|  | 13          | 203.65      | 0.0075   | 0.2443   | 0.2283                                | 0.7067  |
|  | 15          | 190.65      | 0.0024   | 0.1768   | 0.1725                                | 0.6616  |
|  | 17          | 177.65      | 0.0007   | 0.1220   | 0.1277                                | 0.6165  |

# Table 6Schwarzschild runs:1–30 μm with 150,000 and300,000 lines

| Lines Schwarzschild    | 150,000 | 150,000 | 150,000  | 300,000 | 300,000 | 300,000  |
|------------------------|---------|---------|----------|---------|---------|----------|
| K = coefficient        | Surface | 9 km    | 17 km    | Surface | 9 km    | 17 km    |
| %Transparent           | 95.25   | 100     | 100      | 95.25   | 100     | 100      |
| Max K                  | 4596    | 0.690   | 0.001    | 4596    | 0.690   | 0.001    |
| Average K              | 1.418   | 0.0009  | 0.000001 | 1.403   | 0.0009  | 0.000001 |
| K<0.0001               | 23,488  | 136,035 | 149,608  | 46,967  | 272,061 | 299,210  |
| $0.0001 \ge K < 0.001$ | 89,475  | 6730    | 392      | 178,948 | 13,482  | 790      |
| $0.001 \ge K < 0.01$   | 13,792  | 5706    | 0        | 27,554  | 11,399  | 0        |
| $0.01 \ge K < 0.1$     | 8737    | 1204    | 0        | 17,505  | 2408    | 0        |
| $0.1 \ge K < 1$        | 7381    | 325     | 0        | 14,769  | 650     | 0        |
| $1 \ge K < 10$         | 5252    | 0       | 0        | 10,507  | 0       | 0        |
| $10 \ge K < 100$       | 1489    | 0       | 0        | 2983    | 0       | 0        |
| $100 \ge K < 1000$     | 359     | 0       | 0        | 717     | 0       | 0        |
| K>1000                 | 27      | 0       | 0        | 50      | 0       | 0        |
| Total lines            | 150,000 | 150,000 | 150,000  | 300,000 | 300,000 | 300,000  |
|                        |         |         |          |         |         |          |

The solar input at the surface varies with cloud cover, and of course with the four seasons of the year as the Earth traverses its path about the Sun. Progressively warmer days in Northern Hemisphere summers are met with progressively cooler days in the Southern Hemisphere. The temperature of the thermal blanket varies accordingly. Adding passing clouds in some statistical way would lower the above numbers slightly, but not change the conclusion.

One can summarize these calculations as follows: whatever the "*climate-change regime*," whatever surface heat from the Sun on any given day within that regime, that heat is fully absorbed and fully vertically redistributed throughout the troposphere—there is no propensity for  $CO_2$  to store heat in a systematic way over time to produce a climatechange effect (as defined in the introduction).

Why does the integrated effect of  $CO_2$  have so little effect on the total temperature profile? The reason is that the Planck function change with height (temperature) is very strong in reducing the intensity of those relatively few lines with large absorption coefficients. Another reason is that the longwave radiation is diffuse which depletes the intensity rapidly over distance. The diffuse nature of the radiation also leads to the fact that the *net radiation* for a given level (that sent upward at the bottom of a layer, minus that sent downward at the top of a layer) further reduces the adsorbed  $CO_2$  radiation intensity.

Other so-called "greenhouse gases" (some with larger absorption coefficients, but all with significantly less concentration) have their intensity quickly transferred upward and depleted by the same strong Planck function intensity change that applies to  $CO_2$  and  $H_2O$ .

From the historical record and from these calculations one sees that the  $CO_2$  concentration had no impact on temperature. It contributes low-level heating and allows upperlevel cooling for a zero net effect.

Similar calculations (not shown here) with the  $H_2O$  bands in Table 2 provided transparencies at even lower levels than  $CO_2$  as expected from the smaller absorption coefficients. However, the cooling effect of water vapor may be just as great as  $CO_2$  in the *upper levels* (though the concentration of  $H_2O$  is less) because of the water vapor continuum between 8 and 13  $\mu$ m which collectively acts with stronger absorption than the individual lines in that spectral region (Stephens 1994).

## Solar magnetic field/cosmic ray factors affecting *climate change*

The work of (Usoskin et al. 2007) provides a history of Grand Minima since 9500 BC. There have been 27 Grand Minima with various durations (from 30 to 110 years), with various time intervals between events. The summarized result from this work was a weak tendency for Grand Minima to cluster with a quasi-period of about 2400 years, and no clear periodicities were observed. There were not significant *climate-change events* with all of these Grand Minima. They were all tied to a special state of the solar dynamo, and (Charbonneau 2014) has expressed the fact that there is strong intermittency in the solar magnetic activity associated with the dynamo.

A significant improvement in determining which Grand Minima are important for *climate change* came with the work of Sharpe (2008) using Jet Propulsion Laboratory DE405 ephemeris data providing the results in Figs. 10 and 11. His C-14 data from Stuiver et al. (1998). The results confirm the reason for the Medieval Warming and the Little Ice Age (1300–1850) with its three separate Grand Minima (Spörer Maunder, and Dalton).

The synodic period ( $T_{\rm S}$ —two successive conjunctions of the same bodies) of two planets 1 and 2 is given by  $1/T_{\rm S} = 1/T_1 - 1/T_2$  (with  $T_1 < T_2$ ). The sidereal periods for Uranus and Neptune are 84.02 and 164.79 years, respectively. This gives  $T_{\rm UN} = 172$  years. This is the main driver seen in the angular momentum of the Sun about the SSB. The relationship of when a solar Grand Minima occurs always involves these four giant planets in their relationship with the Sun and as depicted in Fig. 11—Uranus, Neptune and Jupiter together and Saturn opposite the Sun. Type A events have a slightly stronger impact, but that detail is not relevant in the present discussion.

Fig. 10 Solar activity from Sharpe (2008) with his C-14 data from INTCAL98 (Stuiver et al. 1998)





Fig.11 Typical planetary positions for important grand minima (Sharpe 2008)

McCracken, et al. (2014) extended Sharpe's results back through the last 9400 years. The data sources for cosmic rays were the cosmogenic radionuclides Be-10 and C-14 from ice core records and tree rings. The data over the entire record confirms Sharp's results, and their statistical analysis found periodicities near 350, 510 and 710 years which closely approximate integer multiples of half the  $T_{\rm UN}$ synodic period:  $T = (T_{\rm UN}/2) N$  years with N = 4, 6, and 8. Using combinations of these periods one could approximate the transition between the various warm and cold periods observed in the past few 1000 years.

The range of the Sun's orbital angular momentum about the SSB varies from near zero to only 25% of the Sun's differential angular momentum driving the solar dynamo (Landscheidt 2003). Thus, the strength of the solar dynamo can outweigh the effect of the Sun/planet positions. Nevertheless, these results over this along period strongly suggest that the solar magnetic field/cosmic ray interaction is the primary cause of major climate-change events over the past 9400 years of the interglacial period.

The 35-year cool period within the current Modern Warming was an example where the Gleissberg cycle imposed only a modest impact on the existing strength of the magnetic field that was in place. The current Modern Warming will continue until the strength of the Sun's magnetic field declines.

Finally, there is the impact of cosmic ray intensity! Massive bright blue stars populate the rotating spiral arms of the Milky Way Galaxy. These massive stars have short lives. Small stars like the Sun live long enough to orbit around the center of the Galaxy many times. The journey about the Milky Way Galactic Core takes approximately 230 million years (Svensmark and Calder 2007). The solar system moves faster than the rotation of the spiral arms; thus, it is repeatedly running through the arms. The spiral arms contain many cosmic rays!

The length of this manuscript does not allow a further summary of that activity, but the reader will find that the Svensmark and Calder (2007) reference will take one over that journey, observing that all the Earth's major ice ages can be attributed to the Sun's magnetic field/cosmic ray interaction.

#### Summary

There is no correlation of  $CO_2$  with temperature in any historical data set that was reviewed. The *climate-change* cooling over the 1940–1975 time period of the Modern Warming period was shown to be influenced by a combination of solar factors.

The cause of the Medieval Warm Period and the Little Ice Age *climate changes* was the solar magnetic field and cosmic ray connection. When the solar magnetic field is strong, it acts as a barrier to cosmic rays entering the Earth's atmosphere, clouds decrease and the Earth warms. Conversely when the solar magnetic field is weak, there is no barrier to cosmic rays—they greatly increase large areas of low-level clouds, increasing the Earth's albedo and the planet cools. The factors that affect these *climate changes* were reviewed in "Solar magnetic field/cosmic ray factors affecting climate change" section.

The calculations of " $H_2O$  and  $CO_2$  in the radiation package" section revealed that there is no net impact of  $CO_2$  on the net heating of the atmosphere. The received heat is simply redistributed within the atmospheric column. This result is consistent and explains the lack of  $CO_2$  correlations with observations in the past.

The current Modern Warming will continue until the solar magnetic field decreases in strength. If one adds the 350-year cycle from the McCracken result to the center of the Maunder Minimum which was centered in 1680, one would have a Grand Minimum centered in the year 2030.

Size constraints limited this review to a proper finish in providing more details about the climate theory of Svensmark, in particular, about the details of cloud formation and the precise timing of the ice ages. However, the reader can profit by reading (Svensmark and Calder 2007).

 $CO_2$  is a valuable asset: providing the input to the plant world for the food all creatures require, and providing fresh oxygen for every breath inhaled by animals and mankind. Hundreds of articles reveal the benefits of increased  $CO_2$ under various environmental conditions. Examples include net productivity with greater water-use efficiency (Kimball 1983), and greater productivity even under conditions of chilling stress (Schwanz and Pelle 2001). Any greater  $CO_2$ resource in the future will prove valuable when the Earth cools again.

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